

A PATH-INTEGRAL QUANTIZATION OF THE STRAIGHT-LINE STRING

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A path-integral quantization of the relativistic straight-line string is proposed. In an explicitly covariant form starting with the initial four-dimensional dynamics of the relative coordinate τ_μ the three-dimensional one is derived. A connection between constraints appearing in the canonical formalism and the path integral quantization is discussed.

1. *Introduction.* Various kinds of string-like models have been employed to describe the world of hadrons ¹. The massless relativistic straight-line string is usually considered as the simplest dynamical basis of the model of hadrons. This model has been quantized in the canonical quantization formalism ². Dynamics of quark and gluon fields leads us to the study of dynamics of the relativistic string with masses at the ends ³, in which the quarks carry a finite fraction of the energy-momentum of the hadron. But even the simplest version of the straight-line string with masses at the ends cannot be made tractable in the canonical quantization formalism ⁴. In this paper we propose a path-integral Lorentz-covariant approach to the quantization of the massless relativistic straight-line string which can be generalized to the case of the straight-line string with masses at the ends ⁶.

2. *Gaussian representation for the action of the straight-line string.* The standard form of the action of the straight-line string in Euclidean space is

$$S = \sigma_0 \int_0^1 d\gamma \int_0^1 d\beta [\dot{w}^2 \cdot w'^2 - (\dot{w} \cdot w')]^{1/2} \quad (1)$$

where $w_\mu(\gamma, \beta)$ are the coordinates of the string world surface

$$w_\mu(\gamma, \beta) = z_\mu(\gamma) \cdot \beta + \bar{z}_\mu(\gamma) \cdot (1 - \beta) \quad (2)$$

and $z_\mu(\gamma), \bar{z}_\mu(\gamma)$ are the coordinates of the ends of the string. The dot and the prime stand for the derivatives over γ and β parameters. Therefore we are to consider and make it tractable the following functional integral in the Euclidean space

$$G = \int Dz_\mu(\gamma) D\bar{z}_\mu(\gamma) \exp[-S]. \quad (3)$$

The action is invariant under the reparametrization

$$\gamma \rightarrow f(\gamma, \beta), \quad w_\mu(\gamma, \beta) \rightarrow w_\mu(f(\gamma, \beta), \beta) \quad (4)$$

with the function $f(\gamma, \beta)$ satisfying the conditions

$$f(0, \beta) = 0, \quad f(1, \beta) = 1, \quad \frac{\partial f(\gamma, \beta)}{\partial \gamma} > 0. \quad (5)$$

It is convenient to introduce a "center of mass" coordinate $R_\mu(\gamma)$ and a relative coordinate $r_\mu(\gamma)$ as follows ³

$$R_\mu(\gamma) = \frac{1}{2}(\bar{z}_\mu(\gamma) + z_\mu(\gamma)) , r_\mu(\gamma) = z_\mu(\gamma) - \bar{z}_\mu(\gamma) \quad (6)$$

so that

$$w_\mu = R_\mu + (\beta - 1/2) \cdot r_\mu. \quad (7)$$

The boundary conditions can be imposed in the Lorentz-invariant way

$$R_\mu(1) - R_\mu(0) = T \cdot u_\mu, u_\mu \cdot u^\mu = 1 \quad (8)$$

and $r_\mu(0), r_\mu(1)$ are also fixed.

To develop a procedure to evaluate path-integral (3) we use the auxiliary fields formalism, as is usually done in the string theory ⁵.

Let us rewrite (3) as

$$\begin{aligned} G &= \int Dr DR Dh_{ab} \exp[-\sigma_0 \int \sqrt{h} d^2 \xi] \delta(\partial_a w_\mu \partial_b w^\mu - h_{ab}(\xi)) = \\ &= \int Dr DR Dh_{ab} \int_{-i\infty}^{+i\infty} D\lambda^{ab} \exp[-\sigma_0 \int \sqrt{h} d^2 \xi] \times \\ &\times \exp[+ \int \sqrt{h} \lambda^{ab} h_{ab} d^2 \xi] \exp[- \int \sqrt{h} \lambda^{ab} \partial_a w_\mu \partial_b w^\mu d^2 \xi] \end{aligned} \quad (9)$$

where $d^2 \xi = d\gamma d\beta$, $\xi_1 = \gamma$, $\xi_2 = \beta$, $h \equiv \det h$.

It is convenient to decompose ⁵

$$\lambda^{ab}(\xi) = \alpha(\xi) h^{ab}(\xi) + f^{ab}(\xi) \quad (10)$$

with

$$f^{ab} h_{ab} = 0, h^{ab} \equiv (h^{-1})^{ab}. \quad (11)$$

In a similar way as it has been done in the case of Nambu-Goto string ⁵ we can prove that in the continuum limit $\alpha(\xi)$ and $f^{ab}(\xi)$ can be replaced by their mean values

$$\langle \alpha(\xi) \rangle \rightarrow \bar{\alpha}, \quad \langle f^{ab}(\xi) \rangle \rightarrow 0. \quad (12)$$

Equation (12) reflects the fact, that $\alpha(\xi)$ is a scalar, while $f^{ab}(\xi)$ is a traceless tensor.

Using (12) we obtain the following expression for G :

$$G = \int Dr DR Dh_{ab} \exp[-(\sigma_0 - 2\bar{\alpha}) \int \sqrt{h} d^2 \xi] \exp[-\bar{\alpha} \int \sqrt{h} h^{ab} \partial_a w_\mu \partial_b w^\mu d^2 \xi] \quad (13)$$

with the new action which is quadratic in w_μ and contains the new auxiliary fields h_{ab} .

3. *Integration over the auxiliary fields.* The invariance (4) makes it convenient to introduce the new variables $\bar{\nu}(\beta)$, $f(\xi)$, $\eta(\xi)$. Separating out the collective mode $\bar{\nu}(\beta)$ and the field $f(\gamma, \beta)$, satisfying conditions (5), we have

$$\bar{h} \equiv \frac{h}{h_{22}^2} = (T\sigma\bar{\nu}(\beta))^2 \left(\frac{\partial f(\gamma, \beta)}{\partial \gamma} \right)^2. \quad (14)$$

And making a simple rescaling of \hbar_{12} we also introduce the variable $\eta(\gamma, \beta)$ instead of h_{12}

$$\hbar_{12} \equiv \frac{h_{12}}{h_{22}} = \left(\frac{\partial f(\gamma, \beta)}{\partial \gamma} \right) (T\eta(\gamma, \beta)) \quad (15)$$

where T enters boundary condition (8). Taking into account the fact, that

$$Dh_{11} Dh_{22} Dh_{12} = D\hbar D\hbar_{12} h_{22}^2 Dh_{22} \quad (16)$$

and using the well known in the string theory formula ⁵

$$D\hbar \sim \exp\left[-\frac{\text{const}}{\epsilon} \int \sqrt{\hbar} d^2\xi\right] D\tilde{\nu}(\beta) Df(\gamma, \beta) \quad (17)$$

where $1/\epsilon \sim \Lambda$ is the ultraviolet cut-off scale, we arrive at the following expression after changing the integration over $d\gamma$ by $Tdf(\gamma, \beta) \equiv d\tau$ and gaussian integration over $h_{22} \geq 0$

$$G = \int DR_\mu Dr_\mu D\tilde{\nu}(\tau, \beta) d\eta(\tau, \beta) \exp[-A] \quad (18)$$

where

$$A = \int_0^T d\tau \int_0^1 d\beta \frac{1}{2\tilde{\nu}} [\dot{\omega}^2 + (\sigma\tilde{\nu})^2 r^2 - 2\eta(\dot{\omega}r) + \eta^2 r^2] \quad (19)$$

and trivial rescaling $z, \bar{z} \rightarrow (\frac{\sigma}{2\tilde{\alpha}})^{1/2} z, (\frac{\sigma}{2\tilde{\alpha}})^{1/2} \bar{z}$ has been done.

At first we notice that the action doesn't depend on $f(\gamma, \beta)$ which reflects the invariance (4). So that the integral over $Df(\tau, \beta)$ can be factored out and it is equal to the volume of the reparametrization group.

In the standard way ⁵ we have introduced the physical quantity σ , which entered expression (14) and (19)

$$\sigma^2 = \tilde{\alpha}(\sigma_0 - 2\tilde{\alpha} + \frac{\text{const}}{\epsilon}). \quad (20)$$

In the action the function $\eta(\tau, \beta)$ is integrated over β being multiplied by function $\tilde{\nu}(\beta)$. In what follows the integration over $\tilde{\nu}$ will be performed by the steepest descent method and in the extremum $\tilde{\nu}((\beta - \frac{1}{2})) = \tilde{\nu}(-(\beta - \frac{1}{2}))$. So we can consider only the class of functions $\tilde{\nu}((\beta - 1/2)^2)$, which are even functions of $(\beta - 1/2)$.

It is convenient to decompose the functions $\eta(\tau, \beta)$ in orthogonal polynomials $P_n(\beta)$ with weight $\nu(\beta) \equiv 1/\tilde{\nu}(\beta)$

$$\eta(\tau, \beta) = \sum_n P_n(\beta) k_n(\tau) \quad (21)$$

$$\int_0^1 d\beta \nu(\beta) P_n(\beta) P_m(\beta) = \delta_{mn}. \quad (22)$$

After gaussian integration over $R(\gamma)$ with the condition (8)

$$\int DR \rightarrow \int D\dot{R} \int_{-i\infty}^{i\infty} d^4\lambda \exp\left[\int_0^T \lambda^\mu (\dot{R}_\mu - u_\mu) d\tau\right] \quad (23)$$

it is easy to prove that the action can be represented, with λ being rewritten as $i\lambda$, in the following form

$$S = \frac{1}{2} \int_0^T d\tau [a_3 \dot{r}^2 + \sigma^2 a_{-1} r^2 + r^2 \sum_{n=1}^{\infty} k_n^2(\tau) - 2i(a_1)^{-1/2} k_0(\tau)(\lambda r) - 2(a_3)^{1/2} (\dot{r}\tau) k_1(\tau) + \frac{\lambda^2}{a_1} + 2i(\lambda u)] \quad (24)$$

where we have introduced the following notation

$$a_1 = \int_0^1 \nu d\beta, \quad a_3 = \int_0^1 (\beta - 1/2)^2 \nu d\beta, \quad a_{-1} = \int_0^1 \frac{d\beta}{\nu}. \quad (25)$$

The function $k_0(\tau)$ enters only the fourth term in this expression and the integration over $Dk_0(\tau) = \prod_{i=1}^N dk_0(\tau_i)$, with N tending to infinity, gives the factor proportional to the infinite product of δ -functions;

$$\prod_{i=1}^{\infty} \delta(\lambda r(\tau_i)). \quad (26)$$

This means that there is a dynamical condition

$$(\lambda r) = 0. \quad (27)$$

Integrations over $d^4\lambda$ and Dk_n with $n \geq 1$ lead (effectively in the limit $T \rightarrow \infty$) to the following expression (up to a change of the measure)

$$G = \int d\nu(\beta) D\tau_\mu \delta(ru) \exp[-\frac{1}{2} \int_0^T d\tau [a_1 + (\dot{r}^2 - \frac{(r\dot{r})}{r^2}) a_3 + \sigma^2 a_{-1} r^2]] \quad (28)$$

where we have used the fact that the extremum value of λ_μ is

$$\lambda_\mu \sim u_\mu. \quad (29)$$

It is important that there are two constraints

$$(ru) \sim (rP) = 0 \quad (30)$$

$$(rp) = 0 \quad (31)$$

where P_μ is the total momentum of the string and

$$p_\mu = a_3 (\dot{r}_\mu - \frac{(\dot{r}\tau) r_\mu}{r^2}) \quad (32)$$

is the relative momentum of the string.

The first constraint (30) means that only transverse to the total momentum P_μ components of r_μ are responsible for the dynamics of the string. The second one (31) reflects the fact that the action doesn't depend on the components of p_μ longitudinal to r_μ .

Let consider the rest system of the meson $u_\mu = (1, \vec{0})$ and go over from the Euclidean to the Minkovsky space

$$d\tau_E \rightarrow id\tau_M \quad (33)$$

The Hamiltonian corresponding to the action (28) is

$$H(\vec{p}, \vec{r}) = 1/2 \left\{ \frac{1}{a_3} \frac{\hat{L}^2}{\vec{r}^2} + \sigma^2 a_{-1} \vec{r}^2 + a_1 \right\} \quad (34)$$

where $\hat{L} = (\vec{r} \times \vec{p})$ is the operator of the angular momentum.

Since the hamiltonian does not contain the radial part of the kinetic term the field \vec{r}^2 is not a dynamical one. Thus in the spirit of the canonical formalism we must exclude the field \vec{r}^2 . This can be done by solving its equation of motion for a fixed value of the orbital momentum l

$$-\frac{l(l+1)}{a_3 \vec{r}^4} + \sigma^2 a_{-1} = 0. \quad (35)$$

Inserting this extremum value of \vec{r}^2 into expression (34) we arrive at the final expression for the hamiltonian

$$H(\nu, l) = \frac{1}{2} a_1 + \sigma \sqrt{a_{-1}/a_3} \sqrt{l(l+1)}. \quad (36)$$

Solving eq.(36) for the extremum of $\nu(\beta)$ with the conditions (25) we find

$$\nu(\beta) = \left(\frac{8\sigma \sqrt{(l+1)l}}{\pi} \right)^{1/2} \frac{1}{\sqrt{1 - 4(\beta - 1/2)^2}} \quad (37)$$

with $\nu(\beta)$ playing the role of the energy density of the string. This solution corresponds to the spectrum of the hamiltonian

$$E_l^2 = M_l^2 = 2\pi\sigma \sqrt{l(l+1)} \quad (38)$$

which agrees with the result obtained for the straight-line string in the canonical quantization formalism ².

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